PHYSICS COLLOQUIUM

Thursday, 23 October 2008
17:00-18:00

3rd Floor Seminar Room

"The Planck Mission"

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Abstract

Planck is an astronomical satellite part of the Scientific Programme of the European Space Agency, due to be launched in early 2009. It is designed to image the anisotropies of the Cosmic Microwave Background (CMB) over the whole sky, with unprecedented sensitivity (\(\Delta T/T \sim 2 \times 10^{-6}\)) and angular resolution (~5 arcminutes). Planck will provide a major source of information relevant to several cosmological and astrophysical issues, such as testing theories of the early universe and the origin of cosmic structure. The ability to measure to high accuracy the angular power spectrum of the CMB fluctuations will allow the determination of fundamental cosmological parameters such as the density parameter (\(\Omega_0\)) and the Hubble constant \(H_0\), with an uncertainty of order a few percent. In addition to the main cosmological goals of the mission, the Planck sky survey will be used to study in detail the very sources of emission which "contaminate" the signal due to the CMB, and will result in a wealth of information on the properties of extragalactic sources, and on the dust and gas in our own galaxy. The ability of Planck to measure polarization across a wide frequency range (30-350 GHz), with high precision and accuracy, and over the whole sky, will provide unique insight into specific cosmological questions, but also into the properties of the interstellar medium. I will present an overview of the Planck mission, its scientific objectives, the key elements of its technical design, and its current status.